# **Advanced Cosmology 2015**

# M2 NPAC: 4 lectures by Hervé Dole **Detailed Plan**

#### Welcome

### **TD 1**

- Exo 1: Why is the night sky dark?
- Exo 2: Computing the density contrast
- Exo 3: Why is there a problem in naïve galaxy formation?

## Introduction slides on cosmology and this course

# Chapter 1: Probes of inhomogeneities: from P(k) to galaxies

- 1. 2-pt correlation function of galaxies
- 2. P(k) and 2-pt correlation function
- 3. Angular correlation function
- 4. Galaxy-matter bias
- 5. BAOs

# **Chapter 2: Cosmic Microwave Background**

- 1. Angular power spectrum on the sphere: Cl
- 2. An estimator of CI gaussianity of DT/T and alms
- 3. Variance and CI and "Cosmic variance"
- 4. CMB Temperature
- 5. Cross-correlations
- 6. What does I mean?

#### Details on CMB, polarization and cosmological parameters

#### Seminar on Planck results

#### **TD 2**

Exo: CMB & BAOs

# **Chapter 3: Introduction to Galaxy Formation**

- 1. Collapsing haloes (Mo & White)
- 2. Cooling of structures
  - a. Dynamical time
  - b. Cooling time
    - i. Definition
    - ii. The cooling function
- 3. Conditions for collapse
  - a. Conditions
  - b. Implications
  - c. Numbers
- 4. t<sub>dyn</sub> vs t<sub>cool</sub>: why galaxies are what they are
- 5. Overcooling problem
- 6. Role and problem of H<sub>2</sub>

Also: visit of the IAS calibration station (Station d'étalonnage)